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- 43. The conversion is valid in soil at depths greater than a few interaction lengths in the lateral direction where the energy spectrum is in equilibrium and is not valid at any depth for pure Fe shielding. For BNL soil ($\rho = 1.8 \text{ g/cm}^3$) L = 53.28 cm.
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- 46. Using the original calibration, a deep penetration lateral shielding measurement in the D-line of the AGS of 126 mrem/hr compares to the CASIM prediction of 65±19 mrem/hr. Details of this comparison are given in a memorandum from A. J. Stevens to D. Beavis dated 12/4/91. Again using this calibration for the measurement reported in Reference 2 predicts 60±5 mrem/hr compared to a measured 48±11 mrem/hr.
- 47. Radiological Control Manual, Brookhaven National Laboratory, 1995.

- 48. Aerial survey of April 3, 1991 prepared by Chas. H. Sells, Inc. The aerial survey map shows 2 ft elevation contours.
- 49. Recall that these estimated dose rates have the neutron quality factor doubled from current standards.
- 50. The attenuation length for soil derived from Figures 2 and 3 is about 67 cm of soil or 120.6 g/cm². This gives a dose reduction factor of about 1.57 per ft of soil. A 14% reduction is derived by replacing 1 ft of soil with 1 ft of soil-equivalent material whose density is 2.4 g/cm³ instead of 1.8 g/cm³.
- 51. "High Occupancy" is defined to be any region occupied more than ½ hour per day (1/16 of an 8 hour day) by a given individual. Occupancy is discussed more fully in Appendix 1.
- 52. A 1/R attenuation is an appropriate short distance extrapolation for a line source of loss which is most often encountered in "beam scraping" losses at accelerators. Combining this fall-off with the nearly point-loss sources assumed is somewhat conservative but is adopted here to compensate for "groundshine" which, although discussed in section 7 of this note, is usually neglected.
- 53. C. Distenfeld and R. Colvett, "Skyshine Considerations for Accelerator Shielding Design," Nucl. Sci. Eng. Vol. 26, p.117, 1966. The expression given in this reference has been multiplied by 2 because skyshine is dominantly low energy neutrons and we are assuming double the current quality factor.
- 54. The equilibrium spectrum used to "calibrate" CASIM (Reference 5) is well represented by the following: $E^{-1.4}$ for E > 7 MeV, $.19887XE^{-.57}$ for 0.1 MeV < E < 7 MeV, $.07389XE^{-1}$ for E < 0.1 MeV.
- 55. The numerical loss values given in sections 3 and 4 have been doubled which corresponds to the total (0.1% for normal losses) loss over the length of the Transfer Line. An Au ion is taken to be equivalent to $197 \times 10.4/28$ protons.
- 56. E. Lessard, private communication. The methodology employed in the text was previously developed by Lessard.

- 57. Evaporation neutrons, of order 1 MeV, are isotropic, while intra-nuclear cascade neutrons, of order 10 MeV, are generally not.
- 58. The induced activity has been estimated to be of the order of 1.5 mrem/hr at 1 ft from the side of the dump following a one hour cooling period and about 10 times this at 1 ft from the downstream end. This estimate will be documented elsewhere.
- 59. In these calculations the (correct) beam elevation shown in Figure 12 implies that the value of R corresponding to 86 ft elevation is 5.1 m. Compare to footnote 5 above.
- 60. The expression for the total neutron flux is the same as that given in section 6 with the exceptions that the correction to 20 MeV is 1.46 for Au ions and the radius is 510 cm.
- 61. The collimator is approximated by a section of steel 25 cm long at a lateral distance >1 cm within a 1.6 cm inner radius vacuum pipe. Interactions are forced to occur uniformly along the length at a 1 mm distance into the steel, i.e., at X = 1.1 cm, Y = 0, 0 < Z < 25 cm. For convenience, the steel (Fe) shielding was assumed between R = 17 cm and R = 47 cm.
- 62. This result follows from considering an Au ion at $10.4 \, \text{GeV/u}$ to be equivalent to $124 \, 28 \, \text{GeV}$ protons. This equivalence follows from comparison of the maximum star densities shown in Figures 2 and 3, and corresponds (approximately) to \sqrt{E} scaling. We have previously assumed (see footnote 20) that global dose, which integrates dose times area, scales linearly with energy. The CASIM calculations indicate that peak dose, at least in the energy regime considered here, scales differently. The version of CASIM used for heavy ion simulation is very slow and the current uncertainty of collimator loss does not justify the effort required for separate calculations for Au and protons.
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